Fusing Reconstruction Algorithm for None-Line-of-Sight Phasor-field Imaging

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Non-line-of-sight (NLOS) imaging allows seeing through the corner by analyzing the scattered light off a relay surface. While Single Photon Avalanche Diode (SPAD) arrays can enable the imaging system to operate in a parallel way and significantly save data collection time, most of the existing demonstrations still use single-pixel SPADs due to arrays' lack of global gates and little fill-in factors, causing low signal-to-noise ratio (SNR). 32 *×* **32 SPAD array has been used to track hidden objects, but there is still a gap in 3D reconstruction using large-scale SPAD arrays. In this paper, we propose an algorithm to fuse the reconstruction results obtained by virtual illuminations of multiple wavelengths, which can improve the phasor-field algorithm's robustness to low SNR measurement. We demonstrate that the proposed algorithm can solve the trade-off between reconstruction SNR and resolution on synthetic datasets.** © 2021 Optical Society of America

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1. INTRODUCTION

 NLOS imaging has been gaining interest as this technology enables the remote access of hidden or unreachable scenes around corners. Therefore, NLOS imaging can be adopted in rescuing scenarios, clinical endoscopic analysis and autonomous vehi-cles applications.

 Single-pixel gated SPADs with a scanning laser spot are cur- rently preferred ways to acquire hidden 3D scenes, with most set-ups using either in single pixel or array format [[1](#page-3-0)]. Accord- ing to existing algorithms, such as back projection [\[2\]](#page-3-1), light- cone transform (LCT) [\[3\]](#page-3-2), f-k migration [[4](#page-3-3)] and virtual wave optics [[5](#page-3-4)], the employment of single-pixel SPAD requires the laser to scan a grid on the relay wall to collect the data. The sub- sequent long acquisition time makes NLOS imaging currently unsuitable for applications outside laboratories [[6](#page-3-5)].

 To imaging moving objects, the acquisition time need to be cut down while maintaining the reconstruction SNR. This im- provement can be realized by: (a) increasing the laser power and repetition rate, (b) adopting SPAD arrays to detect photons

 at multiple points at once, and (c) improving the algorithm's robustness to low SNR data.

 Since (a) can lead to eye safety problems, some studies have been following the idea of (b). Ref.[\[7\]](#page-3-6) provided a phasor-field al-27 gorithm which enables the non-confocal NLOS reconstruction. This algorithm allows collecting photons simultaneously using SPAD, but in experiment they still used a single-pixel SPAD. The main difficulty of using SPAD arrays in NLOS imaging is the small fill-in factor of SPAD arrays (generally 10% for sec- ond generation SPAD arrays, 13.4% for the mentioned latest megapixel SPAD array). Therefore, the active area per pixel is very limited, and each pixel collects light from a very lim- ited area unless the relay surface is far away. When the relay surface is at the same distance with single-pixel SPAD circum- stance, the signal recieved by each array pixel is significantly weaker.

Recently, Renna et al.^{[[8](#page-3-7)]} first demonstrated the possibility of incorporating a SPAD array for NLOS reconstruction using a non-confocal data acquisition scheme combined with a phasor- field method. They used a specially-designed 16 *×* 1 SPAD line array, but their experiment did not get rid of galvo mir- rors and still scanned the laser over a grid of 150 *×* 150 positions. Though large-format time-gated SPAD array with up to 1 Mpixel has been available [\[9\]](#page-3-8), to our knowledge, the largest SPAD array which has been demonstrated in experiments is 32 *×* 32 [\[10](#page-3-9)] due to large arrays' low temporal resolution or detection rate.

 To loosen the restrictions on hardware, we require the re- construction algorithm to be more robust under low SNR cir- cumstances to compensate the SPAD arrays' small fill-in fac- tors, which is what (c) states. Some work has been done to improve the back projection algorithm [\[11](#page-3-10)]. However, back pro- jection algorithm runs more than a hundred times slower than the phasor-field algorithm using Rayleigh Sommerfeld diffrac- tion (RSD). The latter one can reconstruct room-sized scene in seconds, thus more suitable for future real-time imaging.

 In this work, we focus on improving the phasor-field algo- rithm to prepare for the practical application of SPAD arrays as (b) and (c) indicate. To be specific, we look into the theory and properties behind selecting the virtual illumination wavelength for reconstruction. We use the proposed criterion to reconstruct ⁶⁴ the 3D hidden scene with different virtual wavelengths. By fus- ing these results, we use synthetic data to demonstate that it is able to reconstruct the scene when the SNR is extremely low

⁶⁷ (the peak value of photons is below 10) still in seconds.

⁶⁸ **2. PROPOSED FUSING RECONSTRUCTION METHOD**

⁶⁹ **A. Geometry Setup**

⁷⁰ We first illustrate the NLOS measurement setup in Fig.[1](#page-1-0). A sin-

 gle temporal impulse response measurement with illumination α point \mathbf{x}_p and detection point \mathbf{x}_c is denoted as $d(\mathbf{x}_p, \mathbf{x}_c)$, where x_p is the laser illumination point on the relay wall, and x_c is the sensor detection point on the relay wall. Reconstruction of the hidden scene need a set of different measurements. The NLOS measurement can be divided to con-focal and non-confocal by whether x_p and x_c are located at the same position.

⁷⁸ As Fig.[1](#page-1-0) (a) plots, confocal measurement is usually imple-⁷⁹ mented with single-pixel SPADs and beam splitters. This kind ⁸⁰ of setup can result in high SNR ratio especially when applying 81 retroreflective paint. However, this setup requires a mechani-82 cal scanning system to get different $d(\mathbf{x_p}, \mathbf{x_c})$. Fig[.1](#page-1-0) (b) shows 83 the non-confocal measurement where the laser and SPAD sen-84 sor look at different points on the relay wall. With a proper lens ⁸⁵ projecting the SPAD array's pixels to different **xc** on the relay 86 wall, a non-confocal setup allows acquiring different $d(\mathbf{x_p}, \mathbf{x_c})$ 87 at once when fixing. Hence, the application of SPAD arrays re-

⁸⁸ quires the non-confocal measurement.

Fig. 1. NLOS measurement setup. (a) Confocal measurement. (b) Non-confocal measurement.

⁸⁹ Furthermore, when considering the spatial broadening due ⁹⁰ to field of view (FoV), **xp** and **xc** should be 2D Gaussian spots 91 rather than ideal 0-dimensional points. When the distance be-⁹² tween **xp** and **xc** is smaller than the spatial divergence, the 143 93 non-confocal measurement $d(\mathbf{x_p}, \mathbf{x_c})$ should act as the confocal 94 setup. For confocal measurement, we can calculate the intensity ⁹⁵ ratio between the first bounce and third bounce for a specific

⁹⁶ spatial point **xv** of the hidden scene as:

$$
\frac{I_{\text{Third Bounce}}}{I_{\text{First Bounce}}} = \left[\frac{\sigma_x \sigma_y}{2\pi |\mathbf{x_v} - \mathbf{x_c}|^2} \right]^2 \cos \theta_{\mathbf{c}, \mathbf{v}} \cos \theta_{\mathbf{p}} \qquad (1) \underset{\mathbf{p}}{\overset{\mathbf{p}}{\underset{\mathbf{p}}{\mathbf{p}}}} \cos \theta_{\mathbf{p}} \qquad (2) \underset{\mathbf{p}}{\overset{\mathbf{p}}{\underset{\mathbf{p}}{\mathbf{p}}}{\mathbf{p}}}
$$

97 where *c* is the speed of light and σ_x and σ_y are the standard de- viations of the 2D Gaussian function. For room-sized geometry ⁹⁹ and typical sensor parameters, this ratio is around 10^{−7} to 10^{−9}. Since currently available SPAD sensor pixels all go though an avalanch process after a detection event happens, they have a dead period (the typical value is 30 ns to 60 ns) before then

103 can detect the next arriving photon. Thus, the first bounce of 151

¹⁰⁴ high intensity can cause strong pile-up effect and overshadow ¹⁰⁵ the desired third bounce. For most of the commertial SPAD ar-

¹⁰⁶ rays, the pixels do not have a global shutter to gate out the first ¹⁰⁷ bounce simultaneously. Thus, the laser point **xp** should be out-¹⁰⁸ side the total FoV of the SPAD array to avoid pile-up.

¹⁰⁹ **B. Virtual Wavelength Selection for Fusing Reconstruction**

 The phasor field algorithm uses a virtual illumination function which is similar to the exposure function in correlation-based Time-of-Flight (C-ToF) imaging[[12\]](#page-3-11). A C-ToF imaging system illuminates the scene with temporally modulated light source. The radiant intensity is modulated by a sinusoid of a specific fre- quency. In NLOS circumstances, the illumination is convolved 116 with a Gaussian pulse to perform temporal focusing[[13\]](#page-3-12). We have the implemented virtual illumination function:

$$
\mathcal{P}\left(\mathbf{x}_{\mathrm{p}},t\right) = e^{j\omega_{0}t}\delta\left(\mathbf{x}_{\mathrm{p}}-\mathbf{x}_{\mathrm{ls}}\right)e^{-\frac{(t-t_{0})^{2}}{2\sigma^{2}}}
$$
\n(2)

118 where λ_0 is the given reconstruction wavelength and ω_0 = ¹¹⁹ $2\pi c/\lambda_0$, \mathbf{x}_{ls} is the virtual light source position, and t_0 is the fo c cused time. In general, x_{ls} is one of the laser illumination points 121 on the relay wall, and t_0 corresponds to the imaging depth of ¹²² the hidden scene.

BS $\frac{1}{x}$ as wavefront on the relay wall as: ¹²³ When the laser shoots at **x**p and the SPAD pixel points at **x**c, 124 the measured temporal data of arriving photons are denoted ¹²⁵ as *H* ($\mathbf{x}_p \to \mathbf{x}_c$, *t*). To perform the reconstruction, the RSD algorithm picks a certain illumination function $P(\mathbf{x}_p, t)$ first, and ¹²⁷ convolve it by the measured data to get the virtual captured

$$
\mathcal{P}\left(\mathbf{x}_{c},t\right)=\int_{\mathbb{P}}\left[\mathcal{P}\left(\mathbf{x}_{p},t\right)*H\left(\mathbf{x}_{p}\rightarrow\mathbf{x}_{c},t\right)\right]\mathrm{d}\mathbf{x}_{p} \tag{3}
$$

¹²⁹ where the *∗* operator indicates convolution in time.

Object plane ¹³³ length should be able to bring better reconstruction results. Laser SPAD X_V $_{134}$ However, existed literature [[5](#page-3-4)] chooses proper reconstruction ¹³⁰ According to Rayleigh criterion, the spatial resolution is $\Delta = 0.61 \lambda L/d$ where *L* is the imaging distance and *d* is the ¹³² virtual aperture diameter. This criterion shows shorter wavefrequency manually for each measured dataset, the only restric-136 tion is *λ* $≥$ 2∆*x* where ∆*x* is the distance between detection points \mathbf{x}_c on the relay wall. In practice, existed work applies points x_c on the relay wall. In practice, existed work applies $\lambda = 4\Delta x$ or $\lambda = 6\Delta x$ to achieve a better SNR. The guideline ¹³⁹ mentioned above only applies to regular sampling grid, and ig-¹⁴⁰ nores the influence of the temporal resolution of the measure-¹⁴¹ ment system, thus not applicable to different experiment se-¹⁴² tups.

The temporal measured impulse response $H(\mathbf{x}_p \to \mathbf{x}_c, t)$ ¹⁴⁴ can be written as:

$$
H\left(\mathbf{x}_{\mathrm{p}} \rightarrow \mathbf{x}_{\mathrm{c}}, t\right) = \left[f(\mathbf{x}_{\mathrm{p}} \rightarrow \mathbf{x}_{\mathrm{c}}, t) * g(t)\right] * \ell(t)
$$

= $f(\mathbf{x}_{\mathrm{p}} \rightarrow \mathbf{x}_{\mathrm{c}}, t) * \left[g(t) * \ell(t)\right]$ (4)

⁴⁵ where $f(\mathbf{x}_p \to \mathbf{x}_c, t)$ is the photon probability density function
⁴⁶ for measurement $d(\mathbf{x}_p \to \mathbf{x}_c)$, $g(t)$ is the gating window pro-46 for measurement *d*(\mathbf{x}_p → \mathbf{x}_c), *g*(*t*) is the gating window pro-
47 file, and *l*(*t*) is the laser impulse profile.We model the *g*(*t*) and file, and $\ell(t)$ is the laser impulse profile. We model the $g(t)$ and $148 \ell(t)$ as gaussian functions, thus the full width at half maximum ¹⁴⁹ (FWHM) of the measurement system can be estimated by:

$$
\text{FWHM} = \sqrt{t_p^2 + t_d^2} \tag{5}
$$

 $\frac{1}{150}$ where t_p is the pulse width of laser, and t_d is the timebin resolution of SPAD.

 Therefore, it is crutial to adjust the reconstruction wave- length for differnet experiment setups to obtain a suitable SNR. 154 To decide the proper wavelength, we first pull Eq[.2](#page-1-1) into Eq.[3](#page-1-2) $_{201}$ and this generates:

$$
\mathcal{P}_{\mathcal{F}}\left(\mathbf{x}_{c},\omega\right)=(2\pi)^{\frac{3}{2}}\sigma\cdot H_{\mathcal{F}}\left(\mathbf{x}_{ls}\to\mathbf{x}_{c},\omega\right)\cdot e^{-\frac{\sigma^{2}}{2}\left(\omega-\omega_{0}\right)^{2}}e^{-i\omega t_{0}}.\tag{6}
$$

Then we pull in Eq[.4,](#page-1-3) and denotes $h(t) * l(t)$ as $e^{-t^2/2\sigma_t^2}$ where σ_t = FWHM/2.355. The actual virtual captured wavefront's ¹⁵⁸ frequency intensity will satisfy:

$$
|\mathcal{P}_{\mathcal{F}}\left(\mathbf{x}_{c},\omega\right)| \propto e^{-\frac{\sigma_{\mathbf{r}}^{2}}{2}\omega^{2}} e^{-\frac{\sigma^{2}}{2}\left(\omega-\omega_{0}\right)^{2}}
$$
\n(7)

¹⁵⁹ where the former term works as an attenuator and the latter ¹⁶⁰ term works as a selector. $λ_0 = 2πc/ω_0$ should be set larger ¹⁶¹ than a threshold so that the central frequency components of ¹⁶² the latter term can be large enough to perform the reconstruc- 163 tion.

 In RSD method, each frequency component is propogated back into the scene seperately to generate the reconstruction, thus the decay of these high frequency components will cause the reconstruction to be blurrier and dimmer. we conclude that the virtual reconstruction wavelength should satisfy that

$$
\lambda_0 > \alpha \cdot (c \cdot \text{FWHM}) \tag{8}
$$

 to obtain suitable SNR. The parameter *α* has relation to do with the exposure time and laser pulse energy, which decides the intrinsic SNR. The typical value of *α* is 2 *∼* 3 when the peak 172 value of third bounce among all array pixels is around 20.

¹⁷³ **C. Fusion Alogorithm**

 When using the SPAD array in NLOS imaging, the data collec- tion time can be saved by parallel measurement, but the detec- tion rate and FoV of SPAD array pixels are limited. Improve-177 ments in the fabricating can improve the detection rate, but for now SPAD arrays suffer from low detection rate which leads to ill-posed SNR. When using single-pixel SPAD for detection, we usually choose the smallest available wavelength to get the best resolution. When the instrinsic SNR goes down in array cir-182 cumstances, Eq.[7](#page-2-0) indicates that longer wavelength (smaller ω_0) should be employed to suppress the noise. However, this sup- pression would cause a worse resoslution according to Rayleigh criterion. 231

 To solve this contradiction, we propose a simple but ef- fecitive algorithm to fusion the reconstruction results of short and long virtual wavelength. We show that this fusion algo- rithm is able to improve both the SNR and the resolution. Let $I(\mathbf{x}_v, \lambda)$ denotes the reconstruction result using virtual wave-191 length $\lambda_0 = \lambda$, the fusioned result can be expressed a weighted summation for different wavelengths' results:

$$
I(\mathbf{x}_{\mathbf{v}}) = \sum_{n=1}^{N} w_n(\mathbf{x}_{\mathbf{v}}) I(\mathbf{x}_{\mathbf{v}}, \lambda_n)
$$
\n(9)

193 where the weighting factor $w_n(\mathbf{x}_v)$ denotes the credibility of 194 voxel $\mathbf{x}_\mathbf{v}$ for wavelength λ_n .

195 Further, since we have concluded in Eq.[8](#page-2-1) that the SNR is roughly propotional to the virtual wavelength we use, we use the deviation from longer wavelength's reconstruction result to measure the credibility:

$$
w_n = \mathcal{A}\left(\left|I(\mathbf{x}_v, \lambda_{n+1}) - I(\mathbf{x}_v, \lambda_n)\right|\right) \tag{10}
$$

where $A(\cdot)$ is a monotonic decreasing function in range [0, 1]. As for $n = N$, we let $w_N = 0$, and only use the longest wavelength's result as a reference to get better SNR without sacrific-²⁰² ing resolution.

 We find that when the deviation is larger, this voxel's value 204 is more likely to be noise. The function $\mathcal{A}(\cdot)$ works as a coun- terpart of activation function in neural networks. If the quality of a voxel supposed to be better than the others, it is more acti- vated in the final results. Through numerical experiments, we find that the sigmoid function can achieve the most smooth and stable result:

$$
\mathcal{A}(x) = 1 - \frac{1}{1 + e^{-x}}.
$$
 (11)

The computation overhead of the fusion algorithm is negligible compared to the overhead of RSD algothims. Therefore, ²¹² the time and memory overhead mainly depends on the num-²¹³ ber *N* of total wavelengths used in reconstruction. In practice, ²¹⁴ we find that $N = 2$ can improve the quality of reconstructed ²¹⁵ volume significantly.

²¹⁶ **3. NUMERICAL RESULTS**

²¹⁷ **A. Simulation Setup**

²¹⁸ We setup the NLOS simulation using the idea of ray tracer, and ²¹⁹ build a light-weight renderer. We first discretize the surface of 220 the hidden object as squares at 3D position x_v with area S_o and 221 surface normal *α*. The photon probability density function $f(t)$ 222 in Eq.[4](#page-1-3) is calculated as:

$$
f(\mathbf{x}_{\mathbf{p}} \to \mathbf{x}_{\mathbf{c}}, t)
$$

= $\frac{I_0}{(2\pi)^3} \int \frac{S_0 \cos \theta_{\mathbf{c}, \mathbf{v}}}{|\mathbf{x}_{\mathbf{v}} - \mathbf{x}_{\mathbf{p}}|^2} \cdot \frac{\sigma_x \sigma_y \cos \theta_{\mathbf{p}, \mathbf{v}}}{|\mathbf{x}_{\mathbf{v}} - \mathbf{x}_{\mathbf{c}}|^2} \cdot \frac{\sigma_x \sigma_y \cos \theta_{\mathbf{c}}}{|\mathbf{x}_{\mathbf{c}} - \mathbf{x}_{\mathbf{s}}|^2} dS_0$ (12)

where the geometry notations are plotted in Fig[.1](#page-1-0) (a), and the constant parameter I_0 is composed of the laser's wavelength and pulse energy and the SPAD pixels' detection rate. I_0 can be easily calibrated through one measurement.

227 Additionally, in experiment the detection efficiency can be ²²⁸ extremely low (less than one third bounce photon per 10k ²²⁹ frames). We also take the poisson distribution of arriving photons into consideration, the final expression of simulated data is given by a poisson random number with mean $H(\mathbf{x}_p \to \mathbf{x}_c, t)$ 232 in Eq. [4](#page-1-3).

We use 100×100 SPAD array with 31.25 ps timebin width, and the pixels are pojected to $1 \text{ m} \times 1 \text{ m}$ grids on the relay wall. ²³⁵ The laser has 50 ps width and 500 kHz repetition rate. The hidden object size is $0.4 \text{ m} \times 0.4 \text{ m}$. The 3D setup of our simulation and the detected data is shown in Fig.[2](#page-2-2). Eq.4.

We use 100×100 SPAD array with 31.25 ps timebin widt

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en object size is $0.4 \text$

Fig. 2. Simulation details. (a) 3D setup of simulation. (b) simulated histogram data.

²³⁸ **B. Verification**

239 In Fig[.3,](#page-3-13) we fix the scene geometry parameters to verify the in-²⁴⁰ fluence of the system's FWHM. We also fix the pulse energy ²⁴¹ to make sure in each measurement the sum of total detected ²⁴² photons is the same. However, the imaging results with the ²⁴³ same wavelength $\lambda_0 = 4$ cm show significant difference in SNR, 244 which proves the conclusion we get in Eq[.8.](#page-2-1)

Fig. 3. Simulation with different experiment setups and the same virtual wavelength $\lambda_0 = 4cm$. (a) Fix laser pulse width *t^p* = 50*ps*, change SPAD's timebin resolution;(b): fix SPAD's timebin resolution $t_d = 31.25 \text{ps}$, change laser pulse width.

 \sum_{245} In Fig[.4](#page-3-14), we set $t_p = 50$ ps and $t_d = 31.25$ ps, then we can ^{2[4](#page-3-14)6} calculate that the proper $λ_0$ is 4 cm according to Eq[.8.](#page-2-1) Fig.4 (b) 247 and (c) proves that short virtual wavelength brings a good reso- 275 248 lution, but the shot noise is affecting the reconstruction quality. 276 ²⁴⁹ By contrast, the background noise get suppresed but the letters ²⁵⁰ are blurred in long virtual wavelength reconstruction. Constrained in the set $t_p = 50 \text{ ps}$ and $t_d = 31.25 \text{ ps}$, then we can resolution $t_d = 31.25 \text{ ps}$, change laser pulse width 50ps, change SPAD's timebin resolution;(b): fix SPAD's in resolution $t_d = 31.25 \text{ ps}$, change la

Fig. 4. Imaging using different methods. Top row: front view and second row: side view. (a) Ground truth. (b)-(c): RSD method using virtual illumination wavelength *λ*=4cm and *λ*=8cm. (d) Proposed fusing reconstruction method.

251 In Fig[.4](#page-3-14) (d), we fusion the results of $\lambda_1 = 4$ cm and $\lambda_2 =$ 8 cm using the aforementioned weighting factor. Implemented on general CPU (@ 3.30 GHz), the computation time of the pro-254 posed fusing reconstruction method is 0.015 s for $150 \times 150 \times 100$ voxels, while the RSD reconstruction takes 8.8s and 4.5s 100 voxels, while the RSD reconstruction takes 8.8 s and 4.5 s each for different wavelengths.

 We use structural similarity (SSIM) to measure the quality of reconstruction. Fig.[5](#page-3-15) shows that our proposal can improve the 259 SSIM obviously. We further show in Fig.⁵ that the fusion result of only two wavelengths we chose can outperform the ones of all of the single virtual wavelength results.

6000 two chosen wavelengths. (a): Front view; (b): side view. ⁸⁰⁰⁰ length and the proposed fusing reconstruction method using ¹⁰⁰⁰⁰ **Fig. 5.** The SSIM of RSD method using different single wave-

0 ²⁶² **4. CONCLUSION**

268 between SNR and resolution, the proposal improves the phasor-¹⁰⁰⁰ 267 of measured data is incredibly low. From the trade-off property ²⁰⁰⁰₂₆₆ is proved to improve the the imaging quality when the SNR ³⁰⁰⁰₂₆₅ proposed criterion for the selection of virtual wavelength, and ⁴⁰⁰⁰ 264 NLOS imaging using SPAD array. This method is based on the ⁵⁰⁰⁰₂₆₃ In conclusion, we develop a fusing reconstruction method for ²⁶⁹ field algorithm without increasing the computational overhead. ²⁷⁰ This algorithm is simple and practical enough to facilitate the ²⁷¹ use of SPAD arrays in NLOS imaging.

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²⁷⁸ **Data Availability Statement.** Data underlying the results pre-

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